Extensions of Maxwell's electrodynamics by axions

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Outline

- One page summary of Maxwell's theory
- Introducing the axion, a CP = -1 dynamical field
- Effective Electro-Magneto Dynamics of axions
- Self-sustaining homogeneous axion-electromagnetic configuration in EEMD *
- \bullet Electromagnetic Radiation damping in axion stars $^{\dagger\ \ddagger}$



$$L = -\frac{1}{4}F_{\mu\nu}F^{\mu\nu} - j_{\mu}A^{\mu}, \qquad F_{\mu\nu} = \partial_{\mu}A_{\nu} - \partial_{\nu}A_{\mu}$$

Non-relativistic decomposition $F^{0i} = -E_i, \qquad F^{ij} = -\epsilon^{ijk}B_k \qquad \mathbf{B} = \nabla \times \mathbf{A}, \qquad \mathbf{E} = -\dot{\mathbf{A}} - \nabla A_0.$

Dual field strength and Bianchi's identity: $F_{d\mu\nu} = \frac{1}{2} \epsilon_{\mu\nu\rho\lambda} F^{\rho\lambda}, \qquad \partial_{\mu} F_{d}^{\mu\nu} = 0$

Homogeneous Maxwell-equations follow: $\nabla {\bf B} = 0, \qquad \nabla \times {\bf E} = - \dot{{\bf B}}.$

Electrodynamics is *CP*-symmetric $L_{\Theta} = \frac{\Theta}{4} F_{\mu\nu} F_{d}^{\mu\nu} = -\Theta \mathbf{EB} \qquad \text{FORBIDDEN!}$ Dynamical equations $(\delta \int d^4 x L / \delta A_{\nu} = 0)$: $\partial_{\mu} F^{\mu\nu} = j^{\nu}, \qquad \nabla \mathbf{E} = \rho, \quad \nabla \times \mathbf{B} - \dot{\mathbf{E}} = \mathbf{j}$

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Why is QCD CP-invariant? (Origin of axion hypothesis)

Non-abelian gluon-dynamics with CP-odd completion

$$L^{QCD} = -rac{1}{4}F^a_{\mu
u}F^{a\mu
u}, \qquad F^a_{\mu
u} = \partial_\mu A^a_
u - \partial_
u A^\mu - gf^{abc}A^b_\mu A^c_
u.$$

$$L_{\Theta}^{QCD} = \Theta \frac{g^2}{64\pi^2} F^a_{\mu\nu} F^{a\mu\nu}_d, \qquad F^a_{d,\mu\nu} = \frac{1}{2} \epsilon_{\mu\nu\rho\sigma} F^{a\rho\sigma}.$$

Unnatural smallness: $|\Theta_{exp}| < 10^{-10}$.

Peccei-Quinn mechanism

 $U_{PQ}(1)$ spontanously broken at scale f_a provides pseudo-Goldstone, pseudoscalar a(x):

$$\Theta \rightarrow \Theta_{eff} = \Theta + \xi \frac{\langle a(x) \rangle}{f_a} \approx 0.$$

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Model: axion-gauge coupling mediated by heavy fermions

Axion-coupling to electromagnetic field:

$$\begin{split} \mathcal{L}_{EM+a} &= \frac{1}{2} \left[(\partial_{\mu} a(x))^2 - m_a^2 a(x)^2 \right] - \frac{1}{4} F_{\mu\nu} F^{\mu\nu} - j_{\mu} A^{\mu} \\ &+ \frac{1}{4} g_{a\gamma\gamma} a(x) F_{\mu\nu}(x) F_d^{\mu\nu}(x). \end{split}$$

$$F_{\mu\nu} = \partial_{[\mu}A_{\nu]} \qquad \qquad \partial_{\mu}F^{\mu\nu} + g_{a\gamma\gamma}\partial_{\mu}a(x)F^{\mu\nu}_{d} = j^{\nu}_{e}$$

 $a\gamma\gamma$ coupling might arise from PQ-charged heavy quarks



Axion electrodynamics I.

$$L = \frac{1}{2} \left(\epsilon \mathbf{E}^2(x) - \frac{1}{\mu} \mathbf{B}^2(x) \right) - g_{a\gamma\gamma} a(x) \mathbf{E}(x) \cdot \mathbf{B}(x)$$
$$+ \frac{1}{2} \left(\dot{a(x)}^2 - (\nabla a(x))^2 - m_a^2 a^2 \right) - j_0 A_0(x) + \mathbf{j}(x) \mathbf{A}(x),$$

Modified Maxwell equations:

$$\nabla \mathbf{E} + g_{a\gamma\gamma} \mathbf{B} \nabla a(x) = \rho_e,$$

$$-\dot{\mathbf{E}} + \nabla \times \mathbf{B} - g_{a\gamma\gamma} (\dot{a} \mathbf{B} + \nabla a(x) \times \mathbf{E}) = \mathbf{j}_e,$$

Axion equation

$$\ddot{a}(\mathbf{x},t) - \nabla^2 a(\mathbf{x},t) + m_a^2 a(\mathbf{x},t) = g_{a\gamma\gamma} \mathbf{E} \cdot \mathbf{B}.$$

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Axion electrodynamics II.

The axion "medium"

$$\begin{split} \mathbf{P} &= g_{a\gamma\gamma} a \mathbf{B}, \qquad \mathbf{D} = \mathbf{E} + \mathbf{P}, \qquad \nabla \mathbf{D} = \rho, \\ \nabla \times \mathbf{H} - \dot{\mathbf{D}} &= \mathbf{j}, \qquad \mathbf{H} = \mathbf{B} - g_{a\gamma\gamma} a \mathbf{E} \end{split}$$

Or effective charge and current density of electrically neutral axion

$$\rho_{axion} = -\nabla \mathbf{P} = -g_{a\gamma\gamma} \mathbf{B} \nabla a,$$
$$\mathbf{j}_{axion} = \dot{\mathbf{P}} + \nabla \times \mathbf{M} = g_{a\gamma\gamma} (\dot{a} \mathbf{B} + \nabla a \times \mathbf{E})$$

Energy balance

$$W_{charged} = \int d^3 x \mathbf{j} \mathbf{E} = \int d^3 x \mathbf{E} (\nabla \times \mathbf{B} - \dot{\mathbf{E}} - \mathbf{j}_{axion})$$
$$= -\int d\mathbf{F} (\mathbf{E} \times \mathbf{B}) - \frac{d}{dt} \frac{1}{2} \int d^3 x (\mathbf{E}^2 + \mathbf{B}^2) - \int d^3 x \mathbf{j}_{axion} \mathbf{E}$$

After rearangement

$$\frac{dE_{EM}}{dt} + \frac{dE_{axion}}{dt} = \int d^3x \mathbf{j} \mathbf{E} - \int d\mathbf{F} (\mathbf{E} \times B)$$

Review: Paul Sikivie, Rev. Mod. Phys. 93 (2020) 15004 = • • =

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Extensions of Maxwell's electrodynamics by axions

Electromagnetic search for axions

Shining through wall observing axions produced in laboratory Review: A. Ringwald, arXiv:2404.09036 Original proposition Ringwald, Phys. Lett. B**569** (2003) 51



Galactic axion-halo:

standing wavelike $(\lambda \sim km)$, density $\rho_a \sim 0.45 \, GeV/cm^3$ $a(t) = \sqrt{\rho_a} \cos(m_a t)/m_a$

Tools of observation:

i) tunable microwave cavity, hitting resonance ($\omega \approx m_a$) (Sikivie, 1983).

ii) dish antennas (Horns et al., 2013)

Static magnetic field \mathbf{B}_0 parallel to a metallic surface cooperates with axion standing wave to produce oscillating electric field strength $\mathbf{E}(t)$ by

$$\mathbf{E} = -g_{a\gamma\gamma}\mathbf{B}_{0}\dot{a}(t) = g_{a\gamma\gamma}\mathbf{B}_{0}\sin(m_{a}t)\sqrt{
ho_{a}}$$



Axion Electro-Magneto Dynamics

Axions interacting with dyonic heavy quarks (Q) (Sokolov and Ringwald, 2022)

Separate magnetic vector potential C^{μ} is necessary in addition to A^{μ} :

$$G^{\mu\nu} = \partial^{\mu}C^{\nu} - \partial^{\nu}C^{\mu}, \qquad G^{\mu\nu}_{d} = \frac{1}{2}\epsilon^{\mu\nu\rho\sigma}G_{\rho\sigma}$$

 $\partial_{\mu}G^{\mu\nu} = j_m^{\nu}, \qquad \partial_{\mu}G_d^{\mu\nu} = 0$ Bianchi-identity,

Enforcing unique electromagnetic field strength in field equations

 $F_d^{\mu\nu}|_{\mathbf{B},\mathbf{E}} = G^{\mu\nu}|_{\mathbf{B},\mathbf{E}} \rightarrow G^{0i} = -B_i, \qquad G^{ij} = \epsilon^{ijk}E_k.$!! Violates Bianchi-identity on magnetic sources!! Remedy: Dirac, Zwanziger, ...

Axion Electro-Magneto Dynamics

Two new axion-photon couplings

$$\Delta L_{aEMD} = -\frac{1}{8}a(x)\left(g_{EE}F_{\mu\nu}F_d^{\mu\nu} + g_{MM}G_{\mu\nu}G_d^{\mu\nu} + 2g_{EM}F_{\mu\nu}G_d^{\mu\nu}\right).$$

$$\partial_{\mu}F^{\mu\nu} + g_{EE}\partial_{\mu}a(x)F^{\mu\nu}_{d} + g_{EM}\partial_{\mu}a(x)G^{\mu\nu}_{d} = j^{\nu}_{e}$$
$$\partial_{\mu}G^{\mu\nu} + g_{MM}\partial_{\mu}a(x)G^{\mu\nu}_{d} + g_{EM}\partial_{\mu}a(x)F^{\mu\nu}_{d} = j^{\nu}_{m}.$$

In terms of field strengths

$$\nabla \mathbf{E} + g_{EE} \mathbf{B} \nabla a(x) - g_{EM} \mathbf{E} \nabla a(x) = \rho_e,$$

$$\nabla \mathbf{B} - g_{MM} \mathbf{E} \nabla a(x) + g_{EM} \mathbf{B} \nabla a(x) = \rho_m,$$

$$-\dot{\mathbf{E}} + \nabla \times \mathbf{B} - g_{EE} (\dot{a} \mathbf{B} + \nabla a(x) \times \mathbf{E}) + g_{EM} (\dot{a} \mathbf{E} - \nabla a(x) \times \mathbf{B}) = \mathbf{j}_e,$$

$$\dot{\mathbf{B}} + \nabla \times \mathbf{E} - g_{MM} (\dot{a} \mathbf{E} - \nabla a(x) \times \mathbf{B}) + g_{EM} (\dot{a} \mathbf{B} + \nabla a(x) \times \mathbf{E}) = -\mathbf{j}_m.$$

Axion field equation

$$\ddot{a}(\mathbf{x},t) - \nabla^2 a(\mathbf{x},t) + m_a^2 a(\mathbf{x},t) = (g_{EE} - g_{MM}) \mathbf{E} \cdot \mathbf{B} - g_{EM} (\mathbf{E}^2 - \mathbf{B}^2).$$

Axion Electro-Magneto Dynamics

Energy-balance equation (Poynting-theorem)

$$-\int d^3x \left[(\mathbf{j}_e + \mathbf{j}_{axion,e}) \cdot \mathbf{E}(t, \mathbf{x}) + (\mathbf{j}_m + \mathbf{j}_{axion,m}) \cdot \mathbf{B}(t, \mathbf{x}) \right]$$
$$= \int d\mathbf{F} \cdot (\mathbf{E} \times \mathbf{B}) + \frac{d}{dt} \int d^3x \frac{1}{2} \left(\mathbf{E}^2(t, \mathbf{x}) + \mathbf{B}^2(t, \mathbf{x}) \right).$$

With the explicit expressions of axionic currents ($\mathbf{j}_e = \mathbf{j}_m = \mathbf{0}$)

$$-\int d^3x \left[\mathbf{j}_{axion,e} \cdot \mathbf{E}(t, \mathbf{x}) + \mathbf{j}_{axion,m} \cdot \mathbf{B}(t, \mathbf{x}) \right]$$
$$= -\int d^3x \dot{a} \left[(g_{aEE} - g_{aMM}) \mathbf{E} \cdot \mathbf{B} - g_{aEM} (\mathbf{E}^2 - \mathbf{B}^2) \right],$$

by the axion equation equals the rate of change of the axion energy

$$-\frac{d}{dt}\int d^3x \frac{1}{2}\left(\dot{a}^2+(\nabla a)^2+m_a^2a^2\right).$$

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Self-sustaining homogeneous axion-electromagnetic field configuration

$$\begin{split} \mathbf{E}(t), \mathbf{B}(t), a(t) \text{ initiated by external uniform magnetic field } \mathbf{B}_{0} \\ \dot{\mathbf{E}}(t) &= -g_{EE}\dot{a}_{0}(t)(\mathbf{B}_{0} + \mathbf{B}(t)) + g_{EM}\dot{a}_{0}(t)\mathbf{E}(t) \\ \dot{\mathbf{B}}(t) &= g_{MM}\dot{a}_{0}(t)\mathbf{E}(t) - g_{EM}\dot{a}_{0}(t)(\mathbf{B}_{0} + \mathbf{B}(t)). \end{split}$$

Ansätze

$$\mathbf{E} = f_E(a_0(t))\mathbf{B}_0, \qquad \mathbf{B} = f_B(a_0(t))\mathbf{B}_0$$

Initial conditions

$$f_E(a_0(t_0) = a_i) = 0,$$
 $f_B(a_0(t_0) = a_i) = 0.$

Solution with arbitrary $a_0(t)$ $[\lambda^2 = g_{EE}g_{MM} - g_{EM}^2 > 0]$

$$f_E(a_0) = -\frac{g_{EE}}{\lambda} \sin[\lambda(a_0(t) - a_i)],$$

$$f_B(a_0) = \cos[\lambda(a_0(t) - a_i)] - 1 - \frac{g_{EM}}{\lambda} \sin[\lambda(a_0(t) - a_i)].$$

Axion equation emerging after the "elimination" of electromagnetic fields

$$\ddot{a}_0(t) + m_a^2 a_0(t) = \\ = [(g_{EE} - g_{MM})f_E(a_0)(f_B(a_0) + 1) - g_{EM}(f_E(a_0)^2 - (f_B(a_0) + 1)^2)]B_0^2$$

completes the stationary $a_0(t)$, $\mathbf{E}(a_0)$, $\mathbf{B}(a_0)$ configuration induced by \mathbf{B}_0 .

Small amplitude motion $|\lambda(a_0(t) - a_i)| \ll 1$ results in a mass-shift of axions:

$$\ddot{a}_0(t) + M_a^2 a_0(t) = 0, \qquad M_a^2 = m_a^2 + [(g_{EE} - g_{MM})g_{EE} + 2g_{EM}^2]B_0^2.$$

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Status of search for AxionLike Particles (ALP)



Charge quantisation argues for coupling hierarchy $g_{MM} >> g_{EM} >> g_{EE}$

- Electrically neutral axion in extended Maxwell's theory carry effective electric charge and current densities
- Electro-magneto dynamics allows construction of larger set of axion-photon coupling
- In electro-magneto dynamics self-sustaining homogeneous axion+electromagnetic configurations appear, (possibly enhancing dish antenna signal)

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- Gravitationally self-bound axion stars emit electromagnetic radiation at slow rate
- Rotation of polarization plane of electromagnetic radiation propagating through inhomogeneous axion medium
- Echo of electromagnetic radiation from axion cloud via induced axion decay into two photons
- Parametric electromagnetic instabilities of axion stars

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EEMD theory of an axion star

Free axions embedded in $\boldsymbol{B}_{0}\text{-induced}$ medium with gravitational interaction

$$U_{eff} = \frac{1}{2} \left(\dot{a}^2 + (\nabla a)^2 + M_a^2 a^2 \right) + U_{grav}$$
$$U_{grav} = -\frac{G_N}{2} \int d^3 x \int d^3 y \frac{\rho_{axion}(\mathbf{x})\rho_{axion}(\mathbf{y})}{|x - y|}$$

Approximate lowest energy configuration searched in the form $\psi(\mathbf{x}, t) = e^{i\mu_g t}\phi(\mathbf{x})$. μ_g : specific gravitational binding energy

$$\begin{aligned} \mathbf{a}(\mathbf{x},t) &= \frac{1}{\sqrt{2M_a}} \left(e^{-iM_a t} \psi(\mathbf{x},t) + e^{iM_a t} \psi^*(\mathbf{x},t) \right), \\ &\int d^3 x |\psi(\mathbf{x},t)|^2 = N_{axion}, \quad \rho_{axion} = (M_a - \mu_g) |\psi(\mathbf{x},t)|^2 \end{aligned}$$

Variational estimate: $\mu_g/M_a pprox 10^2 (G_N M_a^2 N)^2 << 1$ for $N \sim 10^{61}$.

$$a(\mathbf{x},t) = \sqrt{\frac{2}{M_a}} w F(\xi) \cos(M_a t), \qquad w \sim \sqrt{N}, \qquad \xi = \frac{|\mathbf{x}|}{R}$$

Extensions of Maxwell's electrodynamics by axions

EEMD electromagnetic radiation from an axion star

$$\Box A^{\mu} = j^{\mu}_{axion,e}[\phi, \mathbf{B}_{0}], \qquad \Box C^{\mu} = j^{\mu}_{axion,m}[\phi, \mathbf{B}_{0}],$$
$$\mathbf{e} = -\dot{\mathbf{A}} - \nabla A^{0}, \qquad \mathbf{b} = -\dot{\mathbf{C}} - \nabla C^{0}$$

$$\int d^{3}x \left(\mathbf{j}_{axion,e} \cdot \mathbf{e}(t, \mathbf{x}) + \mathbf{j}_{axion,m} \cdot \mathbf{b}(t, \mathbf{x})\right)$$

$$= -\int d^{3}x \int d^{3}x' \frac{1}{|\mathbf{x} - \mathbf{x}'|} \left[\mathbf{j}_{axion,e}(\mathbf{x}, t) \cdot \frac{\partial}{\partial t} \mathbf{j}_{axion,e}(\mathbf{x}', t - |\mathbf{x} - \mathbf{x}'|) + \frac{\partial}{\partial t} \rho_{axion,e}(\mathbf{x}, t) \rho_{axion,e}(\mathbf{x}', t - |\mathbf{x} - \mathbf{x}'|) \right]$$

$$-\int d^{3}x \int d^{3}x' \frac{1}{|\mathbf{x} - \mathbf{x}'|} \left[\mathbf{j}_{axion,m}(\mathbf{x}, t) \cdot \frac{\partial}{\partial t} \mathbf{j}_{axion,m}(\mathbf{x}', t - |\mathbf{x} - \mathbf{x}'|) + \frac{\partial}{\partial t} \rho_{axion,m}(\mathbf{x}, t) \rho_{axion,m}(\mathbf{x}', t - |\mathbf{x} - \mathbf{x}'|) \right]$$

András Patkós Extensions of Maxwell's electrodynamics by axions

EEMD electromagnetic radiation from an axion star

$$\frac{\overline{dE_{ax}}}{dt}^{T} = -\frac{NB_{0}^{2}X^{3}}{C_{2}} \Big[(g_{EE}^{2} + g_{EM}^{2}) I_{mag}^{2} \\ -\frac{1}{3X^{2}} (g_{EE}^{2} - g_{EM}^{2} - 2g_{MM}^{2}) I_{el}^{2} \Big],$$

where $X = M_a R$ gives the size of the star in units of M_a^{-1} .

$$I_{mag} = 4\pi \int d\xi \xi^2 \frac{\sin(X\xi)}{X\xi} F(\xi),$$

$$I_{el} = 4\pi \int d\xi \xi^2 \left(\frac{\sin(X\xi)}{(X\xi)^2} - \frac{\cos(X\xi)}{X\xi}\right) F'(\xi).$$

Reinterpretation of $\frac{\overline{dE_{ax}}}{dt}^{T}$ as $M_a \frac{\overline{dN(t)}}{dt}^{T}$ yields $N(t) \sim t^{-1/5}$.

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